

ISUSPro
The Satlantic ISUS Nitrate Processor
Users Guide

Version 3.0

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1 Introduction

The Satlantic ISUS Nitrate Processor (**ISUSPro**) is a software utility to be used in conjunction with Satlantic's ISUS Nitrate Sensor. ISUSPro uses previously acquired data and provides a graphical representation of the data. Additionally, ISUS data can be reprocessed using different calibration data and processing parameters.

2 Software Installation

ISUSPro is available as a MS-Windows application. It requires MS-Windows NT, 2000, XP or higher operating system to run. A self-installing version of the program is available from Satlantic's website (www.Satlantic.com). In order to execute the program, the MatLab Component Runtime has to be installed. This component is present for all installations of MatLab 7 (R14). For users without a MatLab license, a stand-alone runtime is available from Satlantic.

Before installing ISUSPro, make sure that there is no other version installed. Otherwise, first uninstall the previous version. To install ISUSPro from CD, insert the CD into a CD drive and start the setup program. This can be done in Windows Explorer, or by selecting "**Run...**" in the "**Start**" menu and typing "**D:\ISUSPro_1.2_Installer.exe**". (The CD drive letter may be different from D.) Follow the instructions on the screen. To install ISUSPro from the self-extracting install program, open the file in Windows Explorer or through the Start menu as described above.

3 Processing Data

3.1 Input Requirements

ISUSPro requires an **ISUS data file** as input. This file can be either logged by SatView (*.RAW) or internally by the ISUS instrument (*.DAT).

In order to read these files, **light and dark Telemetry Definition Format files** (*.TDF), which are distributed with the ISUS are needed.

Finally, a calibration **coefficient file** (ISUS*.CAL) is required for data processing.

Enter these input files either by typing the file names into the provided text boxes, or by using the three browse buttons below the boxes. When browsing for the TDF files, enter both light and dark frame files at the same time.

When processing multiple data files, it is recommended them all together with tDF and CAL files into a separate folder. All intermediate files and optional output files can then be found in sub-folders within that folder.

3.2 Preprocessing

Reading ISUS data files is a lengthy procedure for the ISUSPro software. For this reason, intermediate files are automatically generated in the 'colfiles' sub-folder. These files can be deleted at any time. ISUSPro will regenerate them when needed.

3.3 Processing Parameters

Processing is done for one data file at a time. Before starting to process, select the version of the ISUS instrument: If your serial number is less than 100, the instrument is version 1, otherwise it is version 2. ISUS instrument versions differ by the CAL file format they are using.

Depending on the ISUS version, a number of processing parameters can be adjusted.

For version 1 instruments (radio button 1 checked), a wavelength range of 216.5 to 240 nm should be used. If in the "Curve Fitting Setup" of the ISUS instrument, 3 concentrations are specified, the three extinction coefficient box must be checked. If 2 concentrations are specified, the box must remain unchecked. The "Quadratic baseline", "Seawater darks" and "Use dark averages" boxes are provided for troubleshooting purposes only.

The selected calibration file is then used to process the selected data, when the "Process" button is pressed: Normally, the default processing parameters (wavelength range between 216.5 and 240nm, a linear baseline correction, no third extinction coefficient, and shutter darks) are used for the processing

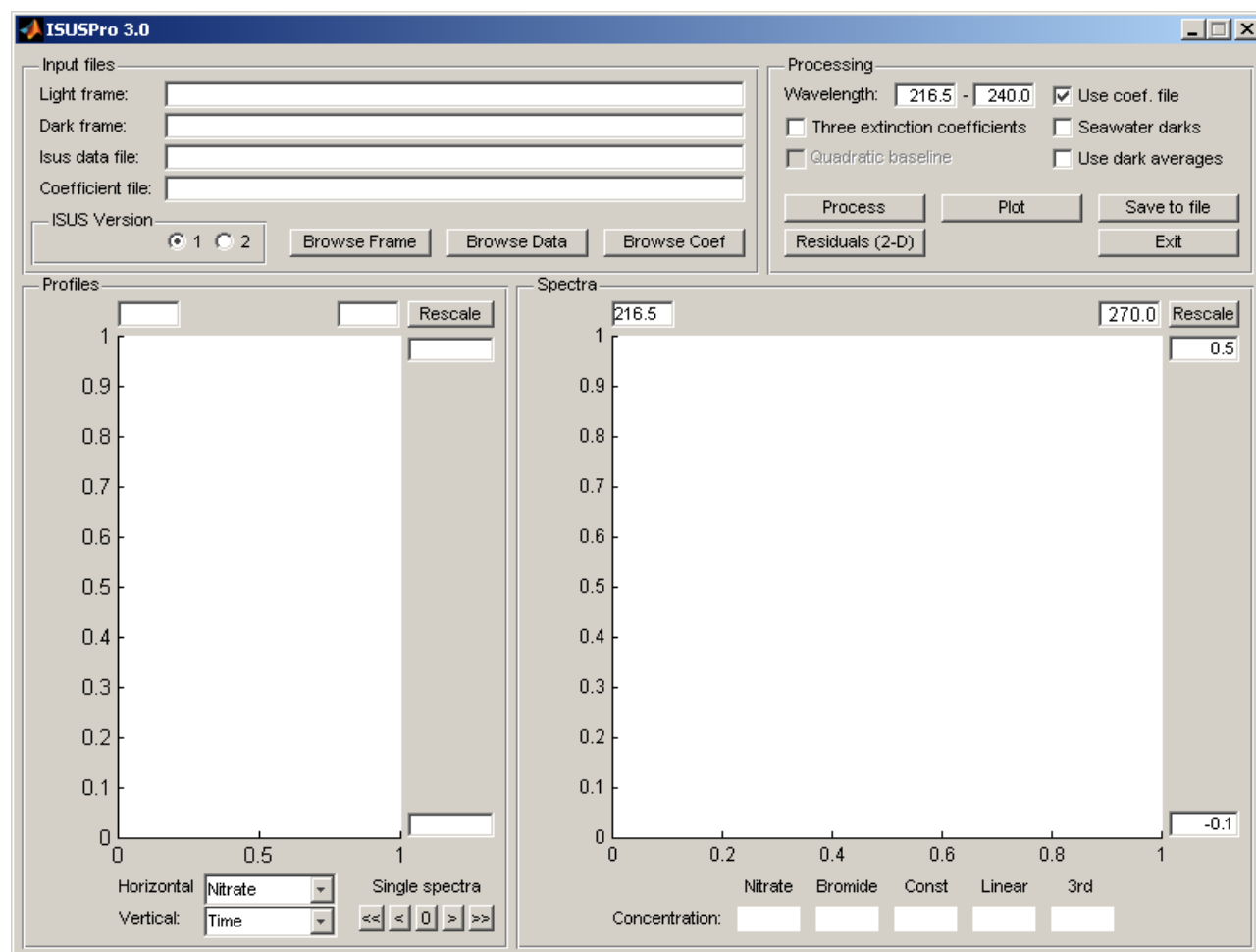


Figure 1: The ISUSPro graphical user interface.

For version 2 instruments (radio button 2 checked), the wavelength range and three extinction coefficient setting is normally taken from the CAL file. In order to overwrite the CAL file settings, the “Use coef. file” box needs to be unchecked.

3.4 Processing

Pressing the “Process” button starts processing. Depending on the size of the data file, processing may take a while. Progress reports are printed in the command window.

4 Viewing Data

Immediately after processing, two plots are displayed (see Figure 2): The left panel contains the nitrate values (horizontal) versus the acquisition time (vertical). The right panel displays averaged spectra and the absorbances of the different chemical species.

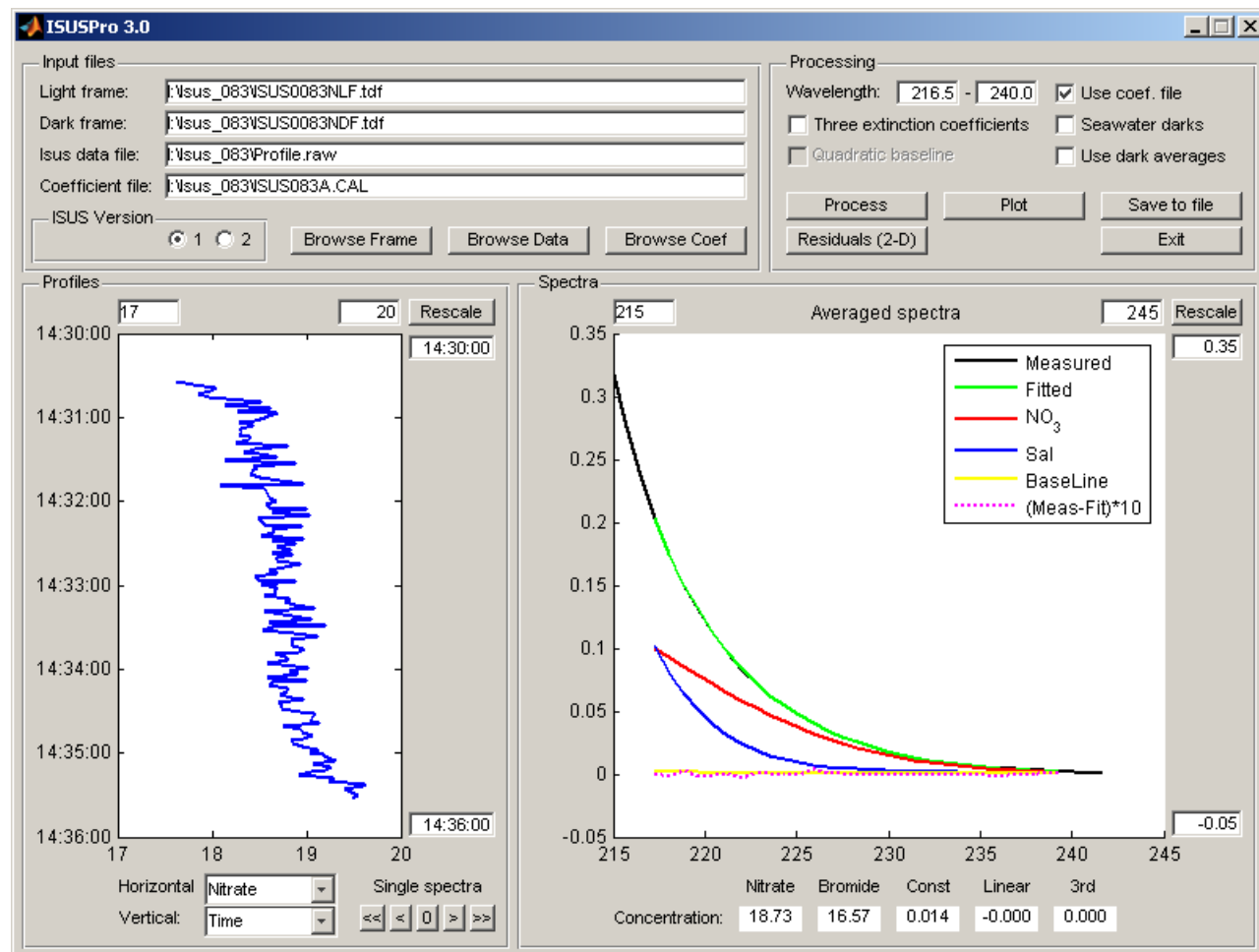


Figure 2: The GUI after processing of the data.

4.1 Profiles

The left panel shows water quality data profiles (horizontal axis) against the acquisition time or the instrument depth (vertical axis). The types of water quality data that can be displayed depend on the instrument setup: Currently supported are fluorescence, CDOM, pH-value, turbidity, temperature, salinity, and oxygen (for ISUS instruments supporting their data acquisition). The vertical axis can display either the time or the instrument depth, depending on whether pressure was recorded.

Two drop-down menus serve to change the data on the horizontal and vertical axis; the change will be effective immediately.

4.2 Spectra

The right panel shows measured absorbance spectra and the components of its deconvolution: baseline, absorbance due to nitrate, bromide (salinity), optional third extinction coefficient, fitted absorbance spectra, and the difference between measured and fitted spectra (multiplied by a factor of 10 for easier viewing). Below the panel, the nitrate value corresponding to the displayed spectrum is displayed.

Immediately after selecting the “Processing” button, an averaged spectrum is displayed. A single spectrum can be selected via the “Single spectra” buttons; a horizontal bar in the profile view indicates which spectra are displayed.

Using the “fast Forward” or “Fast backward” buttons (“<<” or “>>”) causes the software to cycle through all spectra.

4.3 Change of Axis Limits

For both the profile and the spectra panel, the axis limits can be adjusted. For this purpose, there are two input text boxes at each axis of each panel. After entering the values, the respective “Rescale” button will redraw the plot using the new limits.

The vertical axis limits of the left plot show time (hours, minutes, seconds: in HH:MM:SS format) or depth (m). To edit seconds (if a profile is shown as a function of time) one has to first click on the required field using the left mouse button, then move to right using a keyboard button “->”, and finally make the changes as required. Note that even if the starting time in hours is bigger than the end time (e.g. if measurements start at 23:59:00 and end at 00:05:00 of the next day) the indicated time series will be shown correctly since internally a year, month, and day are also considered.

4.4 Residuals

In order to get an overview of the quality of the entire data set, the residuals (i.e. the deviation of the fitted from the measured spectra) can be viewed for all acquired spectra in a 2-dimensional plot. To view the residuals, press the “Residuals (2-D)” button.

4.5 Saving Results

The processed data reside only in the program memory. For future use of processed data, for example within other applications, they can be written to an ASCII file using the “Save tofile” button. The name of the generated ASCII file will be identical to the selected MAT-file. The extension of the ASCII file will be CONC, and the *.CONC files will reside in a sub-folder *asciifiles* within the current data folder. The *.CONC files will contain only those data that were acquired by the ISUS; optional data will not appear via ‘dummy’ columns in the ASCII files.

5 Processing Parameters

The different processing parameters are explained in the ISUS manual and related literature. They are part of the ISSU setup, and it may be desired to 'rerun' an ISUS acquisition using a different set of processing parameters.

The **wavelength range** selects the part of the spectrum to be used for the calculations. All spectral values outside the given wavelength range are ignored. The default range from 216.5nm to 240.0nm was obtained by analyzing data acquired in different scenarios. Special environments may favor a slightly lower or higher range.

The default **baseline model** is linear. This is equivalent to the assumption that any wavelength-dependent instrument drift will be linear. In nearly all situations that were encountered, a linear model was sufficient to account for any drift. Using a quadratic baseline introduces an additional free parameter into the system of equations, and may adversely affect the stability of processing.

Occasionally, ISUS instruments need a **third extinction coefficient** for proper calibration. If this is the case, the calibration file will contain these data. The ISUS-Cal program will add these coefficients during the generation of the cal file if the "use residuals" option is selected. (See the ISUSCal manual for more details.) It is **only in this situation** possible to perform processing with a third extinction coefficient. Normally, the calibration file does not contain a third extinction coefficient. If three extinction coefficients are in use, the ISUS instrument is also set up to utilize three extinction coefficients. This can be confirmed by checking the number of concentrations in the "Curve Fitting Setup" menu of the ISUS instrument.

Dark values of the instrument account for the thermal background noise of the sensors. The dark values are subtracted from the light values to correct for the thermal noise. Normally, the ISUS shutter blocks the path of light at regular intervals to measure the dark values. In the case of shutter failure, the shutter dark values can be replaced by the light values at very short wavelengths using the 'seawater darks' option. At these wavelengths, seawater completely absorbs UV radiation, thus providing an alternate source for the thermal noise correction. **This method does not work in freshwater.** Also, there is a slight difference between using a dark spectrum vs. a dark average when processing data. Assuming there is no shape to the dark spectrum, a dark average should prove more stable; if there is a shape to the dark spectrum, using an average introduces a bias.